

482.

NOTE ON A PAIR OF DIFFERENTIAL EQUATIONS IN THE LUNAR THEORY.

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THE equations

$$\begin{aligned} \frac{d}{dt} \frac{d\rho}{dt} - \rho \left(\frac{dv}{dt} \right)^2 + \frac{1}{\rho^2} &= km^2 \rho \left\{ \frac{1}{2} + \frac{3}{2} \cos(2v - 2mt) \right\}, \\ \frac{d}{dt} \rho^2 \frac{dv}{dt} &= jm^2 \rho^2 \left\{ -\frac{3}{2} \sin(2v - 2mt) \right\}, \end{aligned}$$

taking therein $j = k = 1$ in effect present themselves in the Lunar Theory, and particular integrals in series have been obtained, the development being carried to a great extent; but I give the results only as far as m^4 , viz., writing

$$t - mt = D,$$

we have

$$\begin{aligned} v &= t + \left(\frac{1}{8} m^2 + \frac{5}{12} m^3 + \frac{8}{72} m^4 \right) \sin 2D \\ &\quad + \frac{2}{256} m^4 \sin 4D, \\ \frac{1}{\rho} &= 1 + \frac{1}{6} m^2 - \frac{1}{288} m^4 \\ &\quad + \left(m^2 + \frac{1}{6} m^3 + \frac{1}{18} m^4 \right) \cos 2D \\ &\quad + \frac{7}{5} m^4 \cos 4D. \end{aligned}$$

In the Lunar Theory j and k are properly each $= \frac{1}{1 + \frac{E}{m}}$ (E the mass of the

Earth, m' that of the Sun), but they are taken to be $= 1$; the numerical difference is inappreciable; but there would be a considerable theoretical advantage in retaining

in the equations the coefficients j , k [regarded as each of them = k]: in fact, the developments could then be arranged according to the powers of k , that is according to the powers of the disturbing force; whereas, when k is taken = 1, we have only a development in powers of m , and since m also presents itself through the coefficient $2 - 2m$ of t in $2v - 2mt$, terms which are really of different orders in regard to the disturbing force, are united together into a single term: so that, instead of a term of the form $(Ak + Bk^2 + \&c.) m^p$, where A , B , are numerical, we have the term $(A + B + \dots) m^p$, where of course $A + B + \dots$ is given as a single numerical coefficient. There is no equal advantage in retaining the two coefficients k , j , as this only serves to show how a term arises from the central and tangential forces respectively; thus retaining these coefficients, the integrals as far as m^2 are

$$v = t + \left(\frac{1}{2}k + \frac{7}{8}j\right) m^2 \sin 2D,$$

$$\frac{1}{\rho} = 1 + \frac{1}{8} m^2 k + \left(\frac{1}{2}k + \frac{1}{2}j\right) m^2 \cos 2D,$$

agreeing with the former result when $k = j = 1$; but there is, nevertheless, some interest in retaining the two coefficients. I hope to develop the results somewhat further, and to communicate them to the Society.